

Mapping the Hot Gas in the Universe with the LiteBIRD mission

Mathieu Remazeilles

Instituto de Física de Cantabria (CSIC-UC)

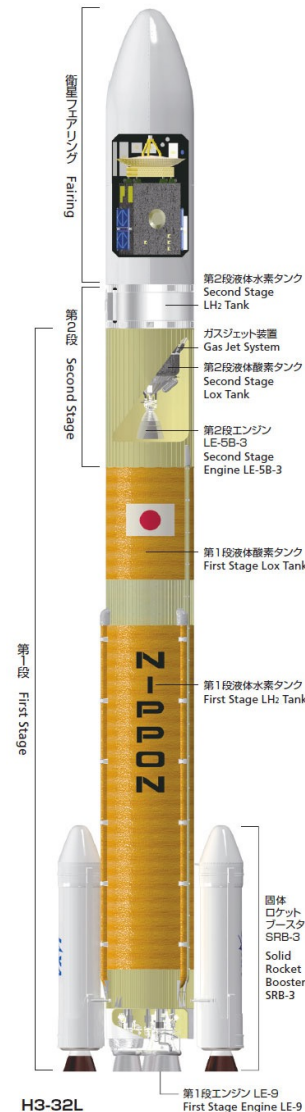
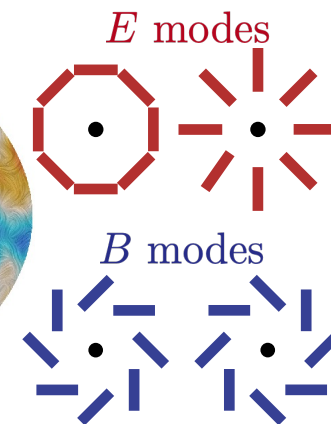
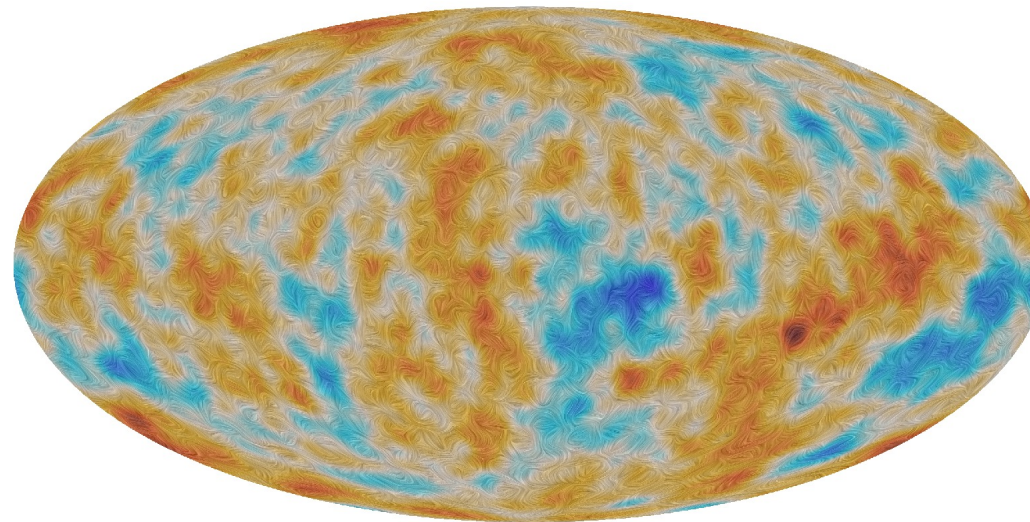
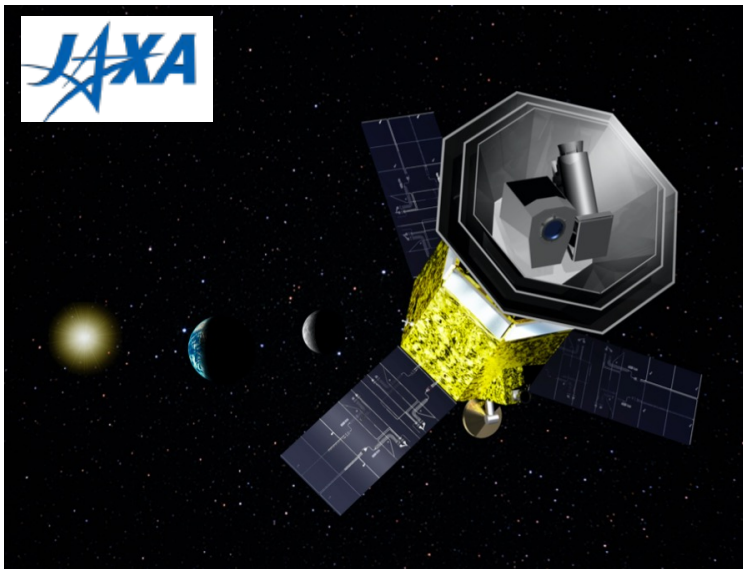


LiteBIRD overview



- Lite (Light) satellite for the study of B -mode polarization and Inflation from cosmic background Radiation Detection
- JAXA's L-class mission selected in May 2019
- To be launched by JAXA's H3 rocket
- **All-sky 3-year survey**, from Sun-Earth Lagrangian point L2
- Large frequency coverage (**40–402 GHz**, 15 bands) at **70–18 arcmin** angular resolution for precision measurements of the **CMB B -modes**
- Final combined sensitivity: **$2.2 \mu\text{K} \cdot \text{arcmin}$**

LiteBIRD collaboration
PTEP 2023





PREPARED FOR SUBMISSION TO JCAP

LiteBIRD Science Goals and Forecasts. Mapping the Hot Gas in the Universe

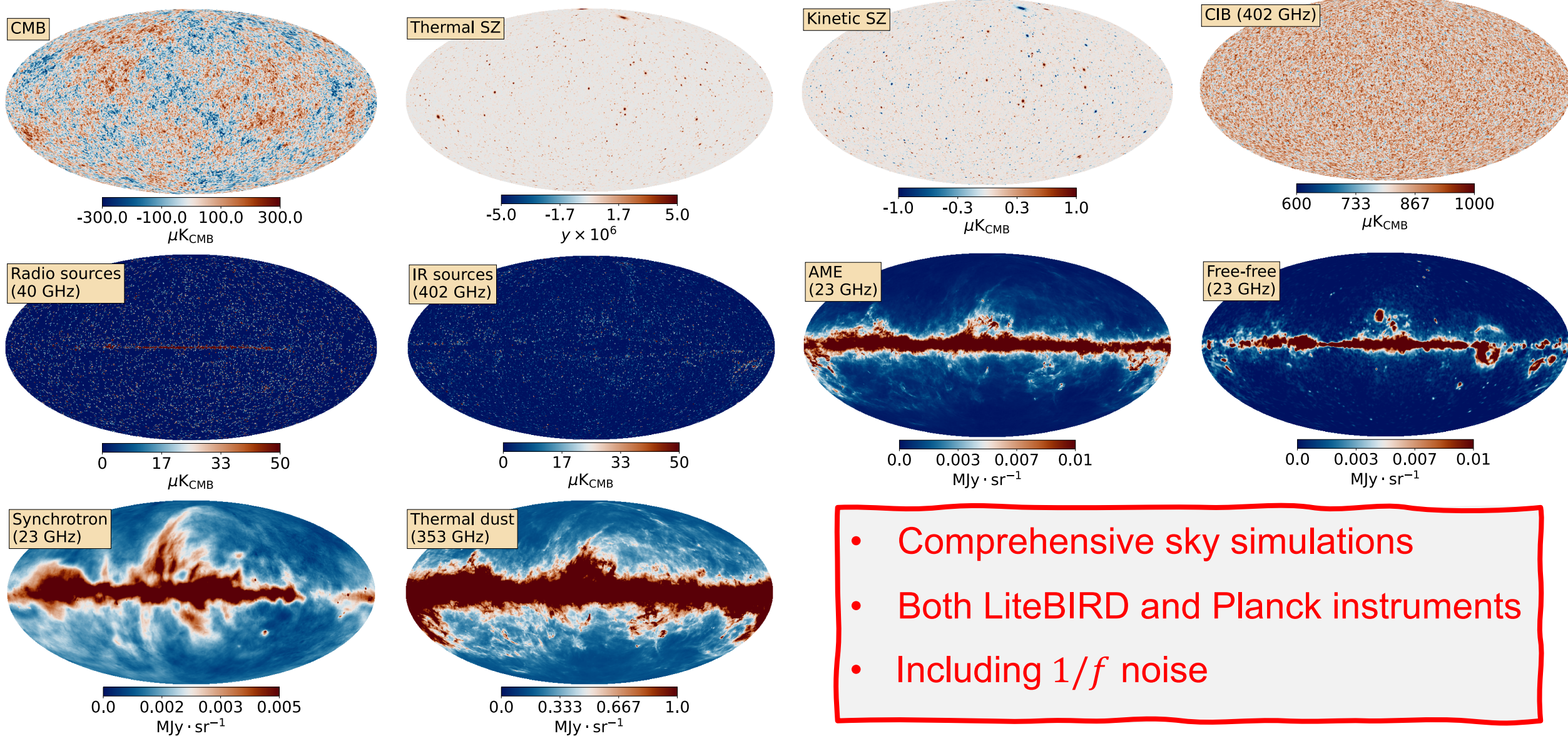
M. Remazeilles,¹ M. Douspis,² J. A. Rubiño-Martín,^{3,4}
A. J. Banday,⁵ J. Chluba,⁶ P. de Bernardis,^{7,8} M. De Petris,^{7,8}
C. Hernández-Monteagudo,³ G. Luzzi,⁹ J. Macías-Pérez,¹⁰
S. Masi,^{7,8} T. Namikawa,¹¹ L. Salvati,² H. Tanimura,¹¹
K. Aizawa,¹² A. Anand,¹³ J. Aumont,⁵ C. Baccigalupi,^{14,15,16}
M. Ballardini,^{17,18,19} R. B. Barreiro,¹ N. Bartolo,^{20,21,22} S. Basak,²³
M. Bersanelli,^{24,25} D. Blinov,^{26,27} M. Bortolami,^{17,18}
T. Brinckmann,¹⁷ E. Calabrese,²⁸ P. Campeti,^{18,29,30} E. Carinos,⁵
A. Carones,¹⁴ F. J. Casas,¹ K. Cheung,^{6,31,32,33} L. Clermont,³⁴
F. Columbro,^{7,8} A. Coppolecchia,^{7,8} F. Cuttaia,¹⁹ T. de Haan,^{35,36}
E. de la Hoz,^{37,1,38} S. Della Torre,³⁹ P. Diego-Palazuelos,^{29,38}
G. D'Alessandro,^{7,8} H. K. Eriksen,⁴⁰ F. Finelli,^{19,41} U. Fuskeland,⁴⁰
G. Galloni,^{17,13} M. Galloway,⁴⁰ M. Gervasi,^{42,39}
R. T. Génova-Santos,^{3,4} T. Ghigna,³⁶ S. Giardiello,²⁸
C. Gimeno-Amo,¹ E. Gjerløw,⁴⁰ R. González González,³
A. Gruppuso,^{19,41} M. Hazumi,^{36,35,43,11,44} S. Henrot-Versillé,⁴⁵
L. T. Hergt,⁴⁶ D. Herranz,¹ K. Kohri,³⁵ E. Komatsu,^{29,11}
L. Lamagna,^{7,8} M. Lattanzi,¹⁸ C. Leloup,¹¹ F. Levrier,⁴⁷
A. I. Lonappan,⁴⁸ M. López-Caniego,^{49,50} B. Maffei,²
E. Martínez-González,¹ S. Matarrese,^{20,21,22,51} T. Matsumura,¹¹
S. Micheli,⁷ M. Migliaccio,^{13,52} M. Monelli,²⁹ L. Montier,⁵
G. Morgante,¹⁹ Y. Nagano,⁵³ R. Nagata,⁴³ A. Novelli,⁷ R. Omae,⁵³
L. Pagano,^{17,18,2} D. Paoletti,^{19,41} V. Pavlidou,^{26,27} F. Piacentini,^{7,8}
M. Pinchera,⁵⁴ G. Polenta,⁹ L. Porcelli,⁵⁵ A. Ritacco,^{52,47}
M. Ruiz-Granda,^{1,38} Y. Sakurai,^{56,11} D. Scott,⁴⁶ M. Shiraishi,⁵⁶
S. L. Stever,^{53,11} R. M. Sullivan,⁴⁶ Y. Takase,⁵³ K. Tassis,^{26,27}
L. Terenzi,¹⁹ M. Tomasi,^{24,25} M. Tristram,⁴⁵ L. Vacher,¹⁴
B. van Tent,⁴⁵ P. Vielva,¹ I. K. Wehus,⁴⁰ B. Westbrook,³¹
G. Weymann-Despres,⁴⁵ E. J. Wollack,⁵⁷ M. Zannoni,^{42,39} and
Y. Zhou³⁶
LiteBIRD Collaboration.

[arXiv:2407.17555](https://arxiv.org/abs/2407.17555)

- Unlike cluster catalogues, which only capture thermal Sunyaev-Zeldovich (SZ) emission from massive, well-resolved clusters, the Compton y-map probes the entire hot gas distribution over the sky.
- The Planck Compton y-map is the first and unique all-sky map of the thermal SZ effect to date.
- Despite low angular resolution for galaxy cluster science, LiteBIRD offers enhanced sensitivity, full-sky coverage, and multiple frequency bands compared to Planck.
- LiteBIRD is well-positioned to deliver the next all-sky thermal SZ map, with reduced foreground contamination compared to Planck.
 - *Important legacy data from LiteBIRD*
 - *Important impact on cosmology and astrophysics*
- We propose to combine both LiteBIRD and Planck channels to leverage the advantages of each experiment for optimal y-map reconstruction and improved constraints on σ_8 .



Sky simulations for LiteBIRD and Planck



- Comprehensive sky simulations
- Both LiteBIRD and Planck instruments
- Including $1/f$ noise

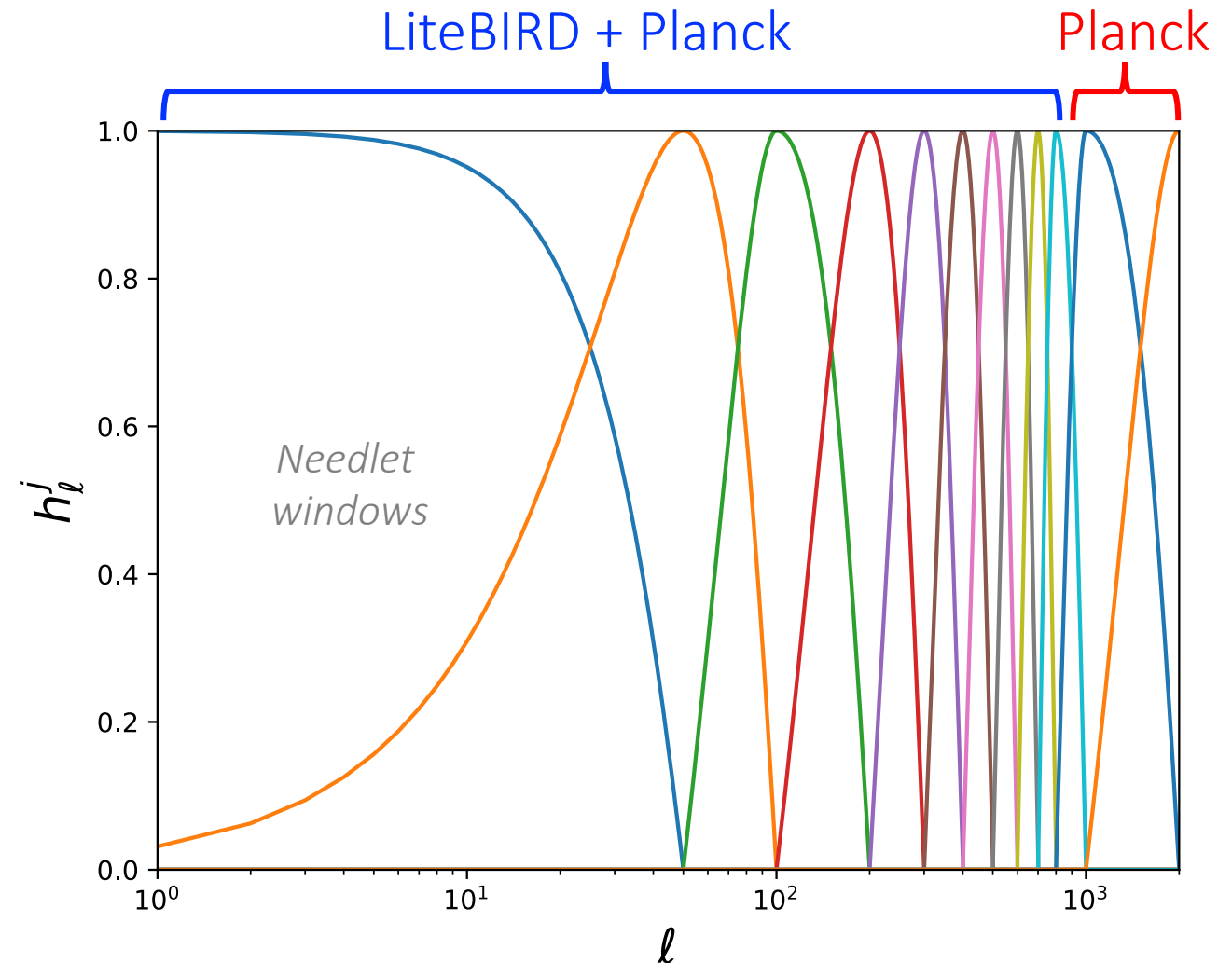


Combining LiteBIRD and Planck channels for component separation with NILC

*NILC enables the combination
of multi-resolution frequency
data from different experiments*

*Remazeilles, Aghanim, Douspis
MNRAS (2013)*

LiteBIRD channels enhance foreground
cleaning, while Planck channels provide
resolution beyond LiteBIRD beam limits



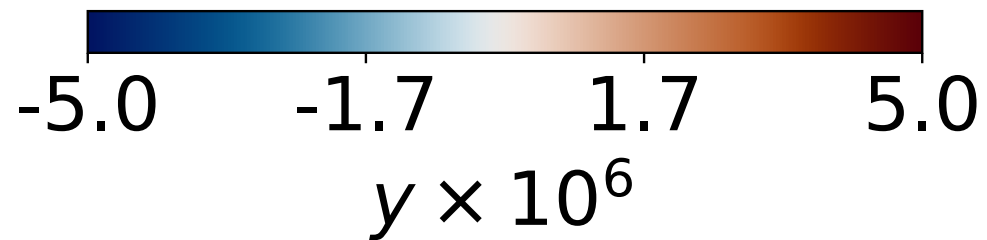
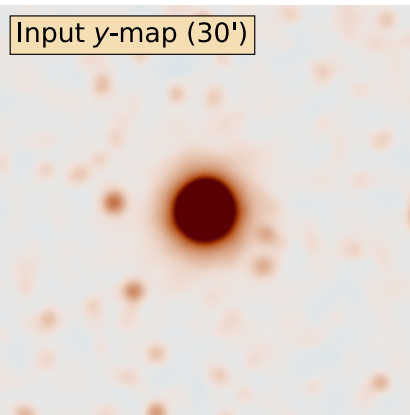


Thermal SZ y-map reconstruction

[arXiv:2407.17555](https://arxiv.org/abs/2407.17555)

Input y-map (30')

red spots are galaxy clusters



$$\langle yy \rangle \sim \sigma_8^8$$
$$\langle yyy \rangle \sim \sigma_8^{11}$$



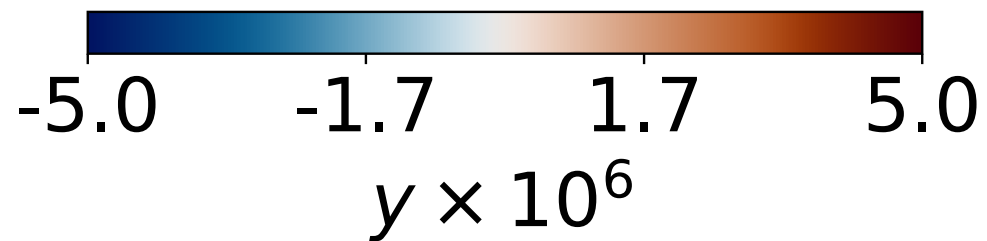
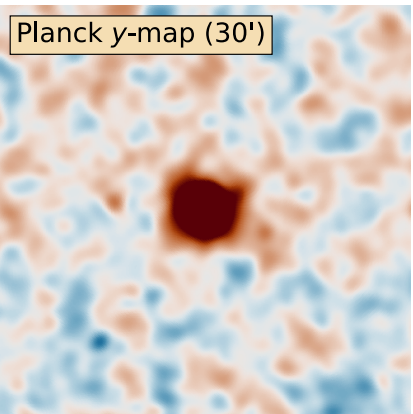
Thermal SZ y-map reconstruction

[arXiv:2407.17555](https://arxiv.org/abs/2407.17555)

Planck
NILC y-map (30')

*After component
separation*

red spots are galaxy clusters



$$\langle yy \rangle \sim \sigma_8^8$$
$$\langle yyy \rangle \sim \sigma_8^{11}$$



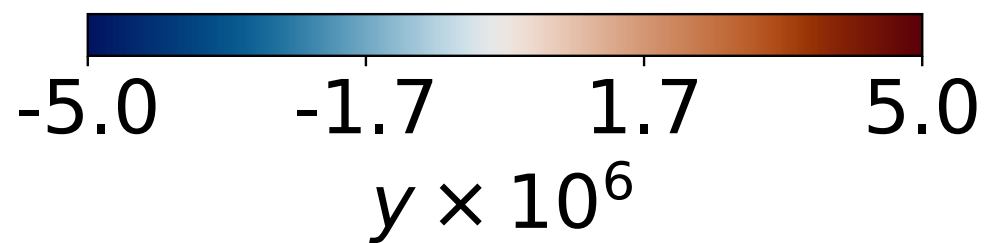
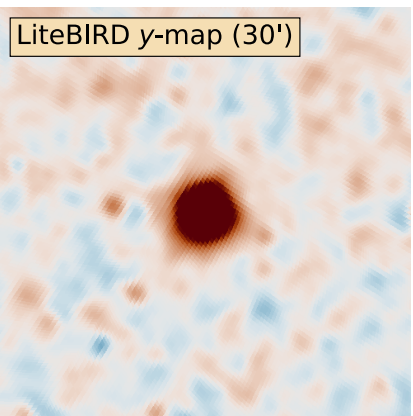
Thermal SZ y-map reconstruction

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LiteBIRD
NILC y-map (30')

*After component
separation*

red spots are galaxy clusters



$$\langle yy \rangle \sim \sigma_8^8$$
$$\langle yyy \rangle \sim \sigma_8^{11}$$



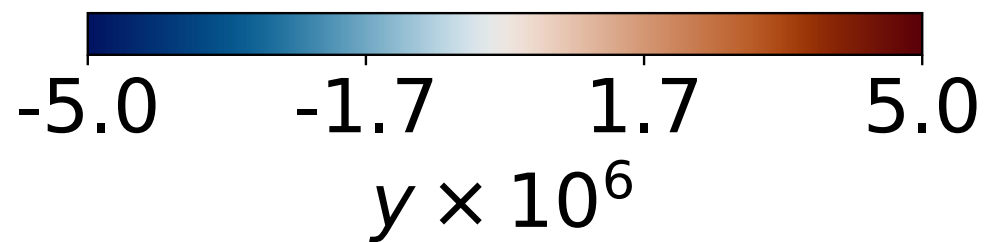
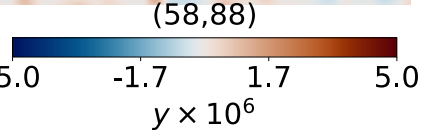
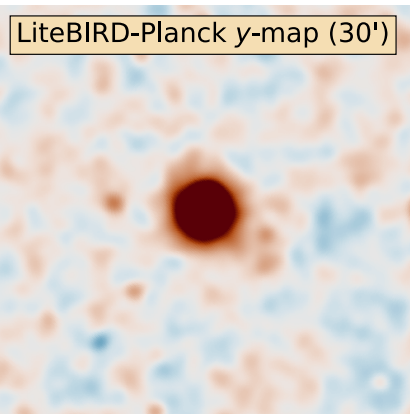
Thermal SZ y-map reconstruction

[arXiv:2407.17555](https://arxiv.org/abs/2407.17555)

LiteBIRD-Planck
NILC y-map (30')

*After component
separation*

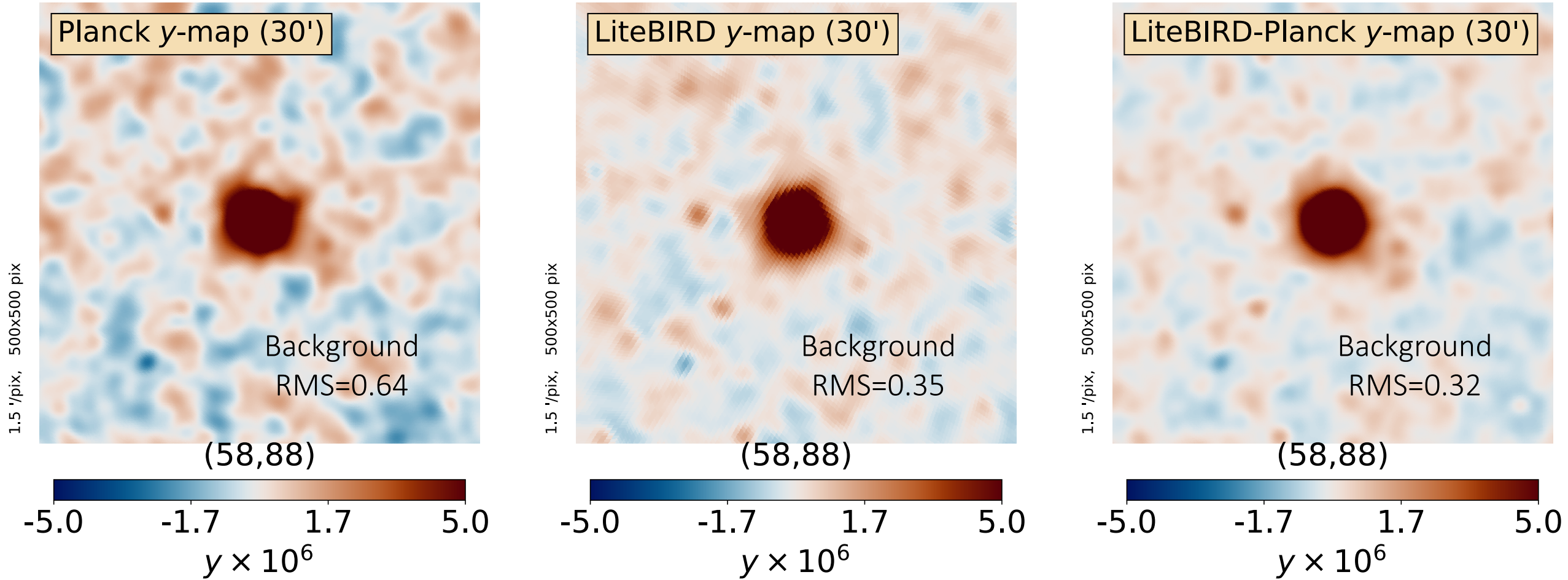
red spots are galaxy clusters



$$\langle yy \rangle \sim \sigma_8^8$$
$$\langle yyy \rangle \sim \sigma_8^{11}$$



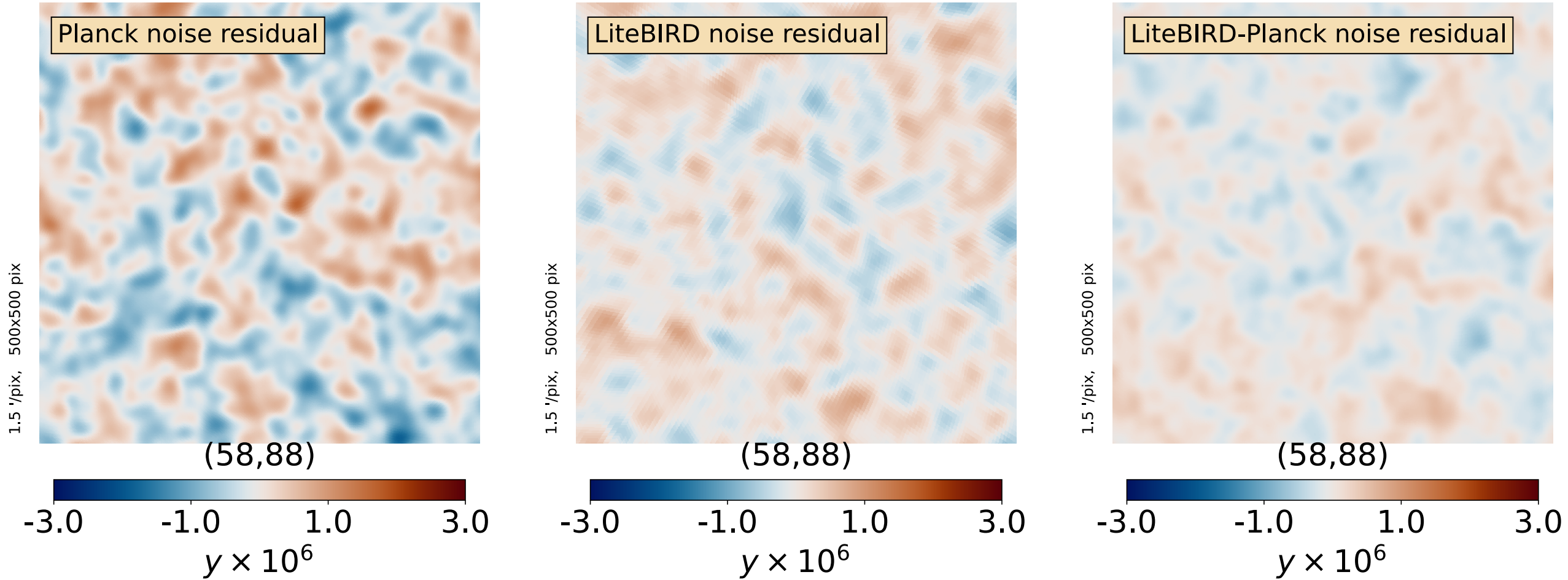
Comparison of y-maps around Coma



*Improvement in SZ map quality from Planck to LiteBIRD,
and from LiteBIRD to joint LiteBIRD-Planck*



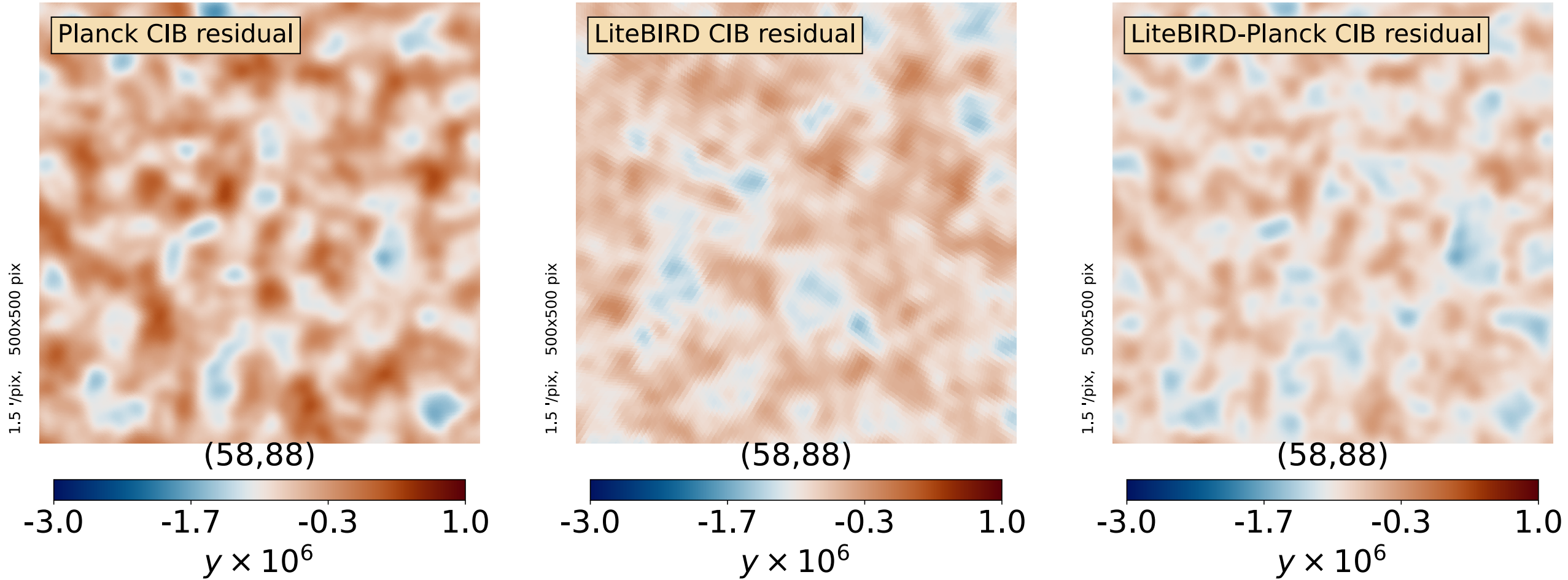
Residual noise contamination



*Reduction in noise contamination from Planck to LiteBIRD,
with further reduction in the joint LiteBIRD-Planck y-map*



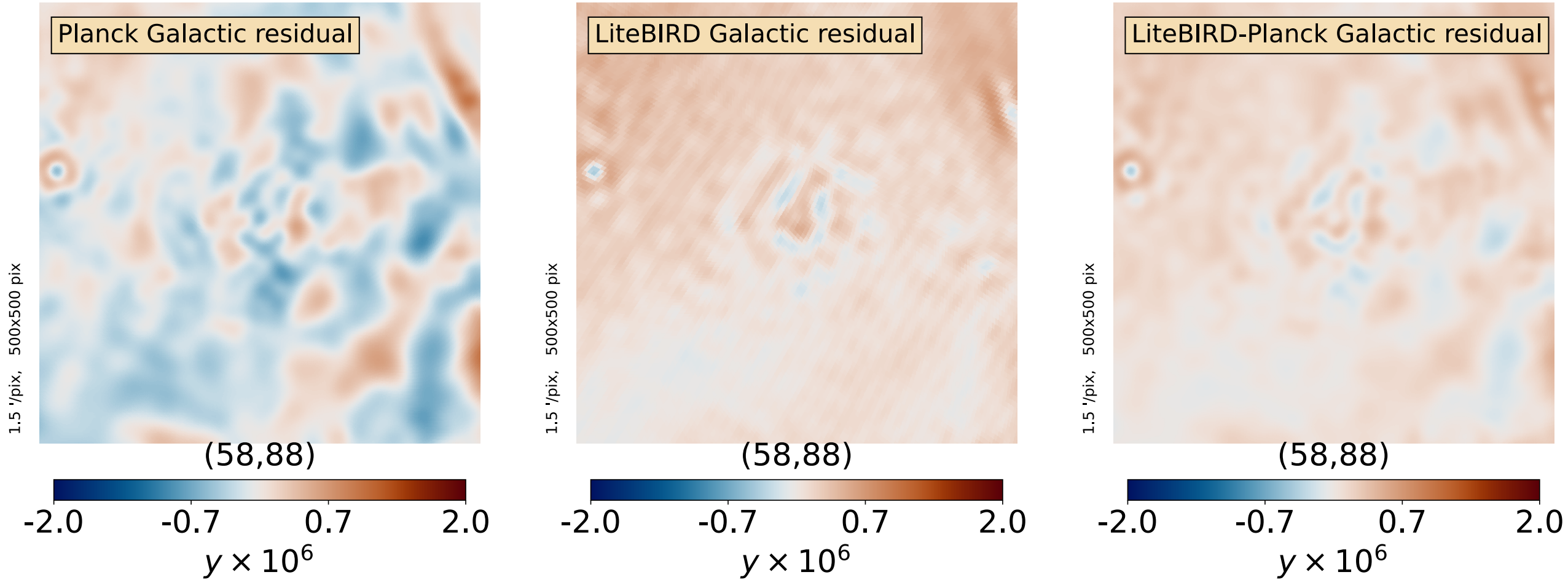
Residual CIB contamination



*Reduction in CIB contamination from Planck to LiteBIRD,
with further reduction in the joint LiteBIRD-Planck y-map*



Residual Galactic foreground contamination



*Reduction in Galactic contamination from Planck to LiteBIRD,
with further reduction in the joint LiteBIRD-Planck y-map*



Reconstructed y -map over 98% of the sky

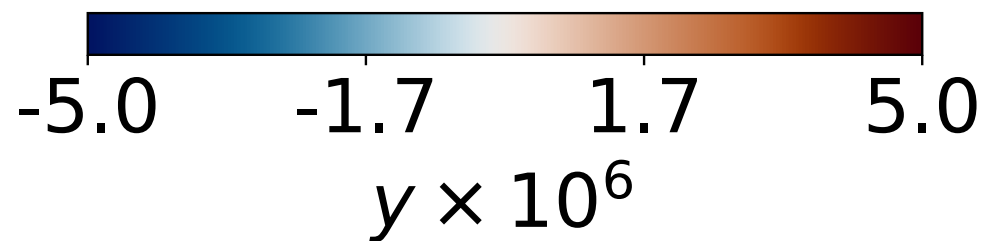
[arXiv:2407.17555](https://arxiv.org/abs/2407.17555)

Planck
NILC y -map (30')

*After component
separation*

98% sky

*Significant contamination
in the Galactic plane*



$$\begin{aligned}\langle yy \rangle &\sim \sigma_8^8 \\ \langle yyy \rangle &\sim \sigma_8^{11}\end{aligned}$$



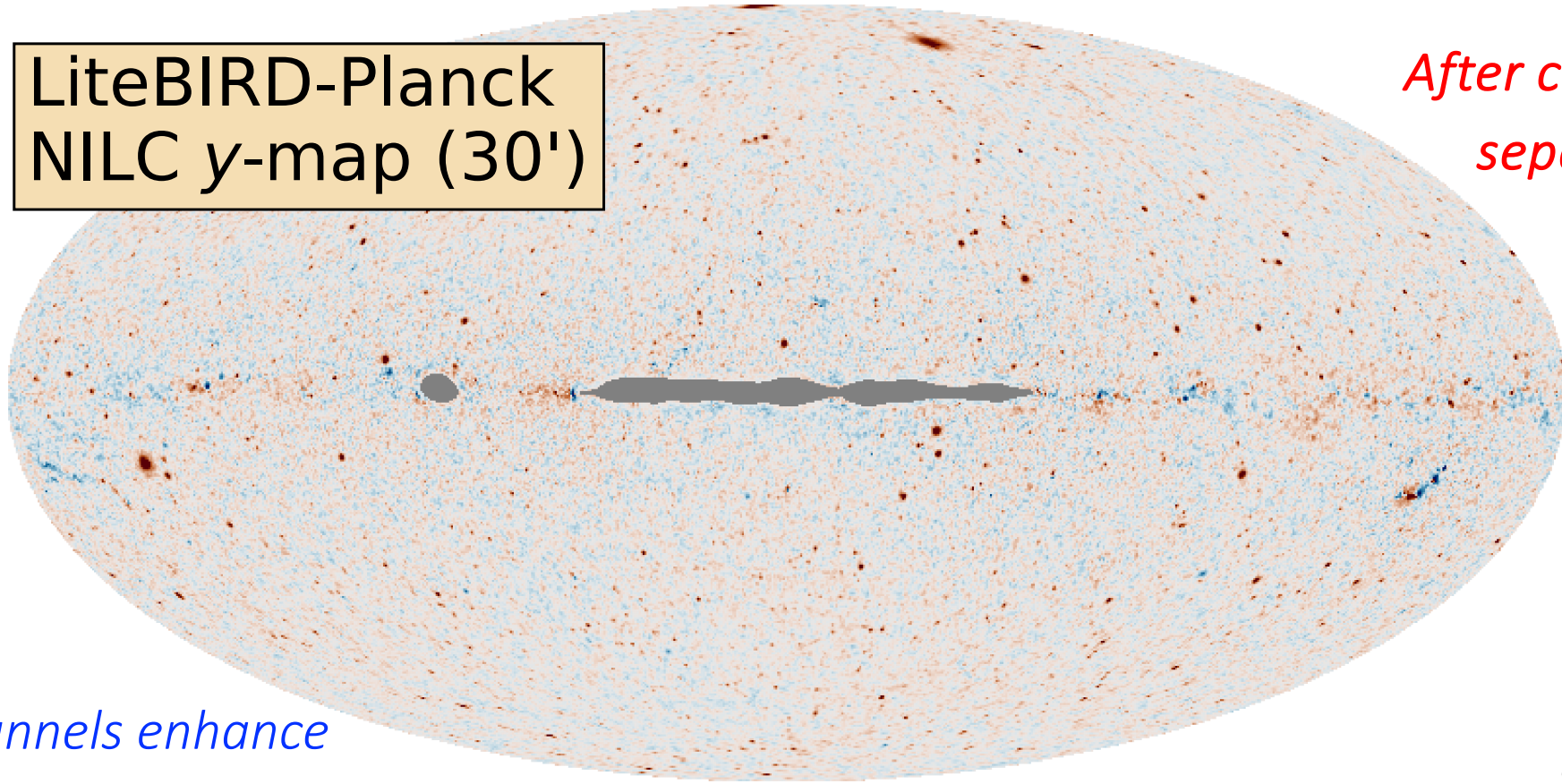
Reconstructed y-map over 98% of the sky

[arXiv:2407.17555](https://arxiv.org/abs/2407.17555)

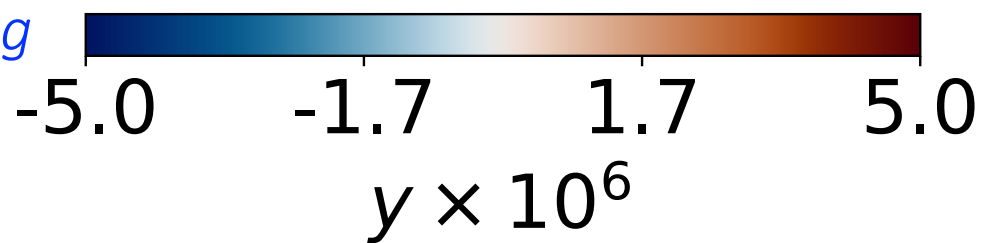
LiteBIRD-Planck
NILC y-map (30')

*After component
separation*

98% sky



*LiteBIRD channels enhance
foreground cleaning, revealing
clusters in Galactic plane*



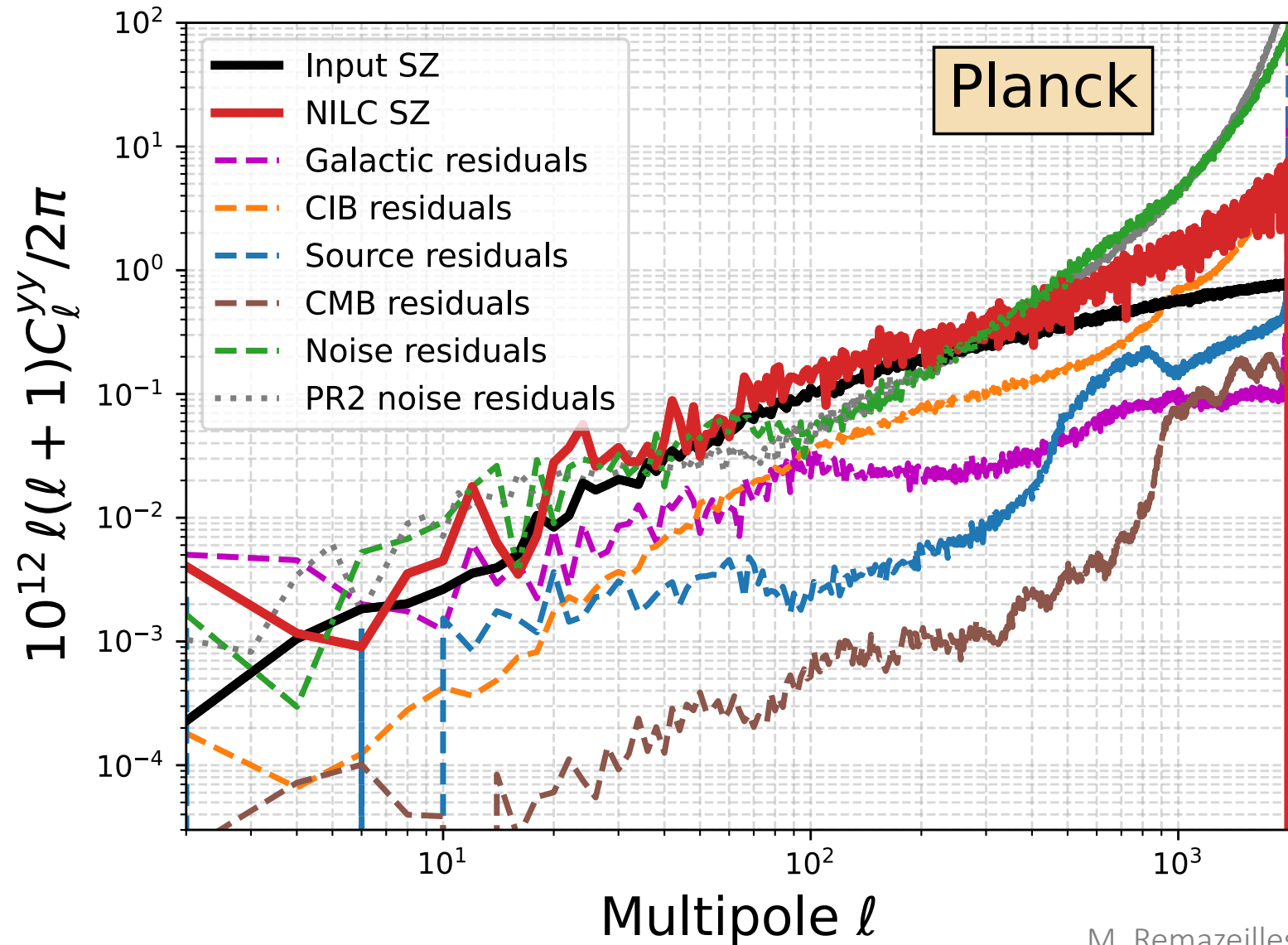
$$\begin{aligned}\langle yy \rangle &\sim \sigma_8^8 \\ \langle yyy \rangle &\sim \sigma_8^{11}\end{aligned}$$



y-map power spectrum and residuals

SZ power spectrum
and residuals

Planck



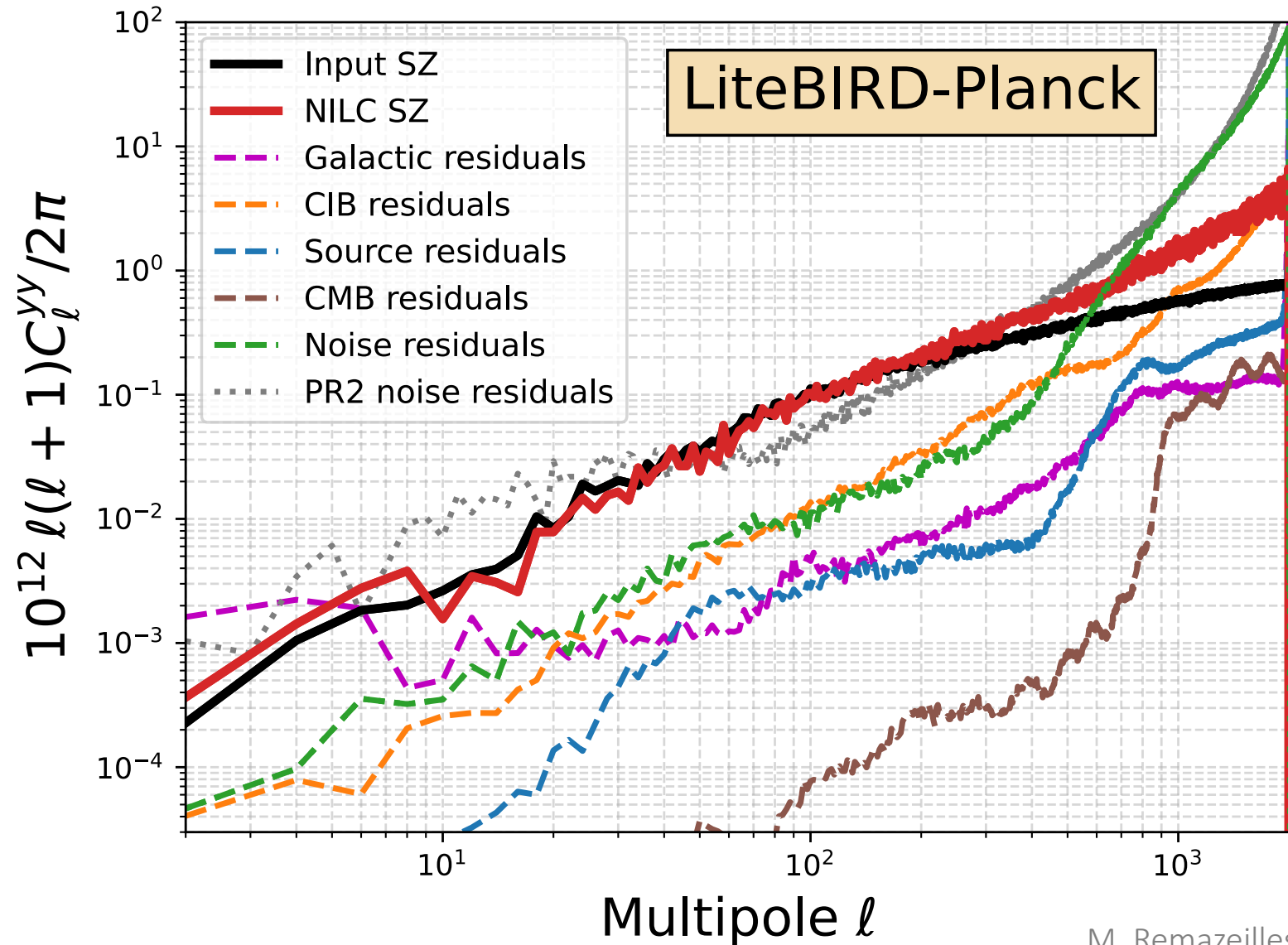


y-map power spectrum and residuals

SZ power spectrum
and residuals

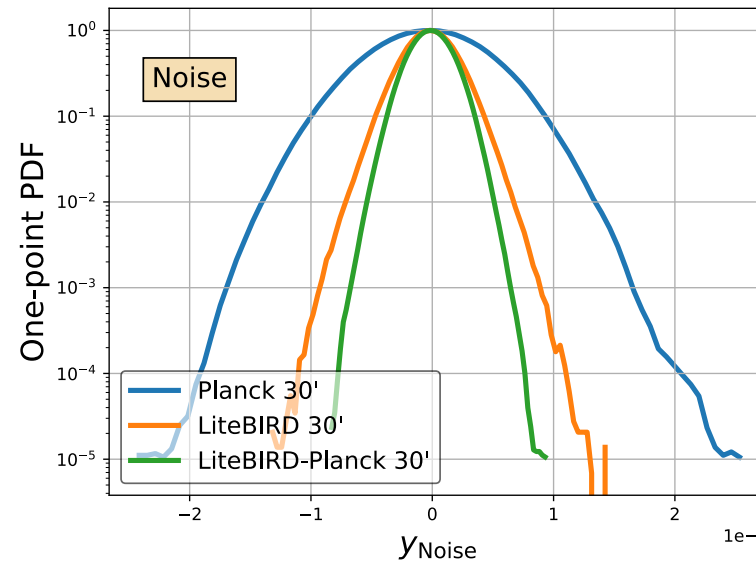
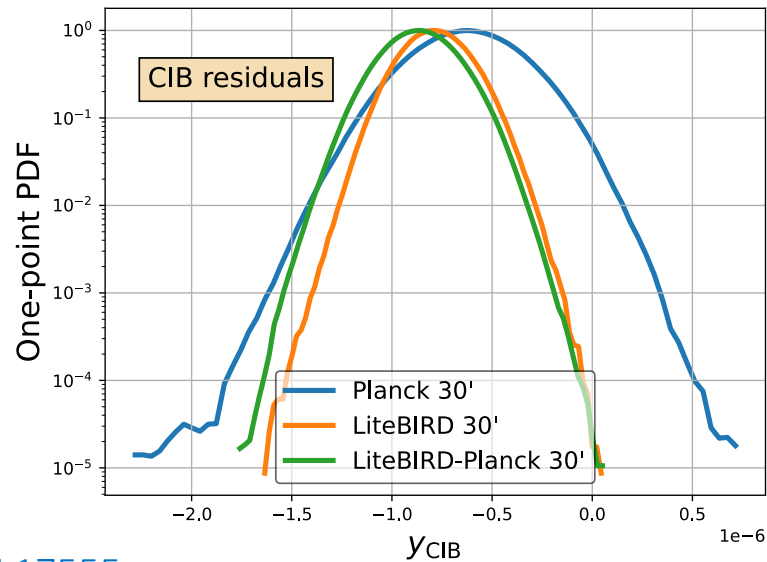
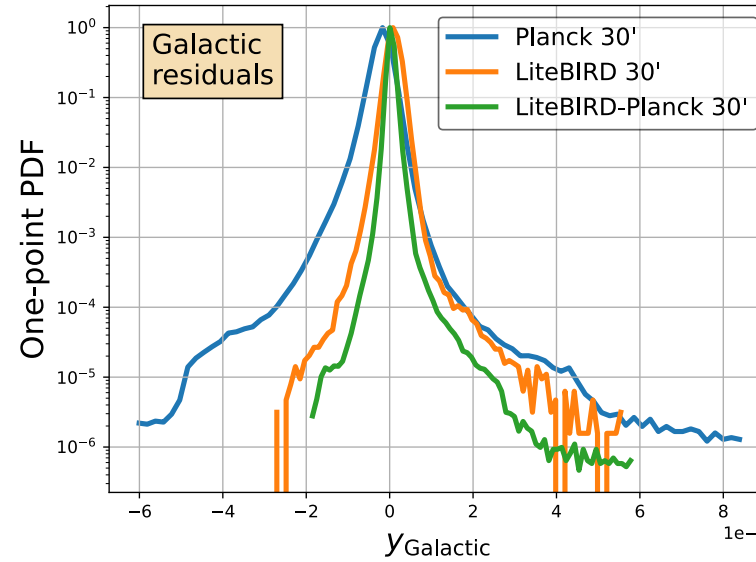
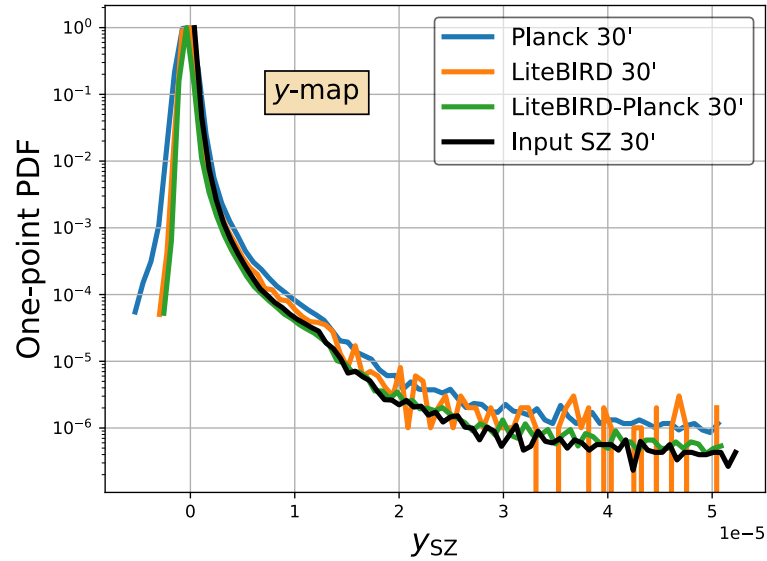
LiteBIRD + Planck

*Noise and foreground residuals
reduced by an order of magnitude
at large and intermediate scales*





One-point PDF of y -map and residuals



Reduction of noise and foreground residuals from Planck to LiteBIRD, with further reduction in joint LiteBIRD-Planck y -map



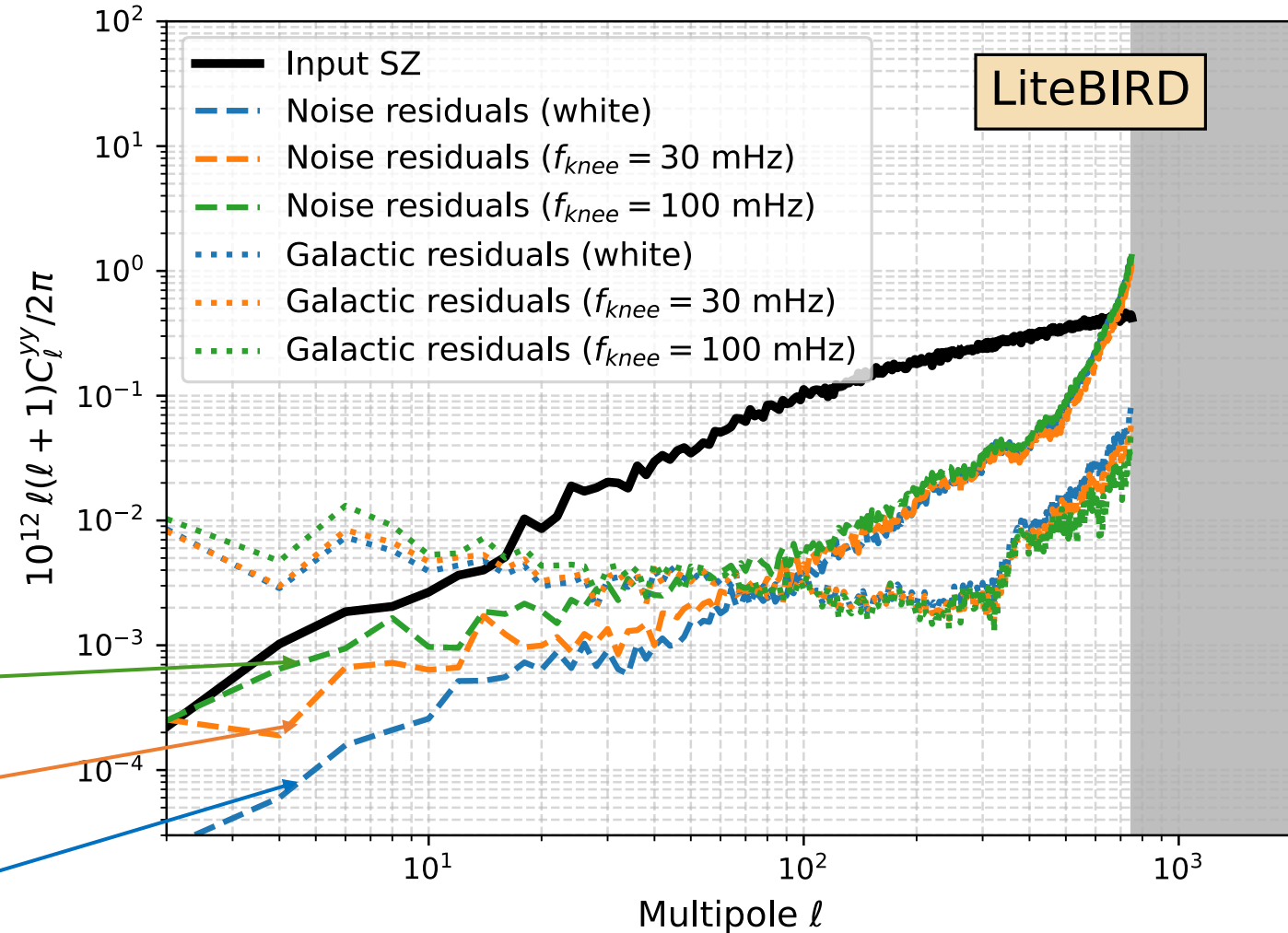
Impact of LiteBIRD $1/f$ noise

LiteBIRD $1/f$ noise reduced below the SZ signal at all multipoles after component separation with NILC

$1/f$ noise ($f_{\text{knee}} = 100$ mHz)

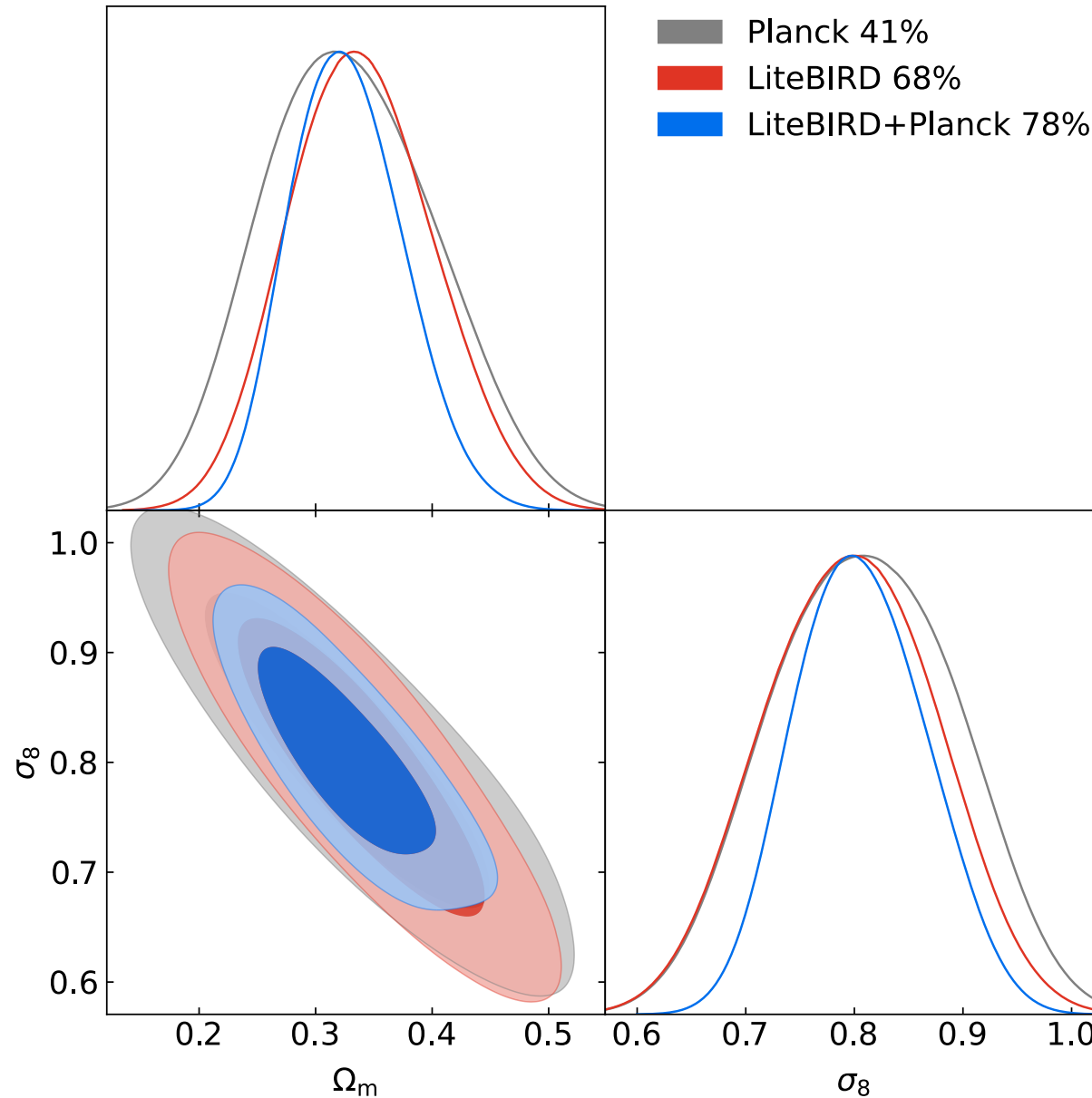
$1/f$ noise ($f_{\text{knee}} = 30$ mHz)

white noise





Cosmological parameter constraints



*Non-Gaussian
contribution to SZ
cosmic variance
included*

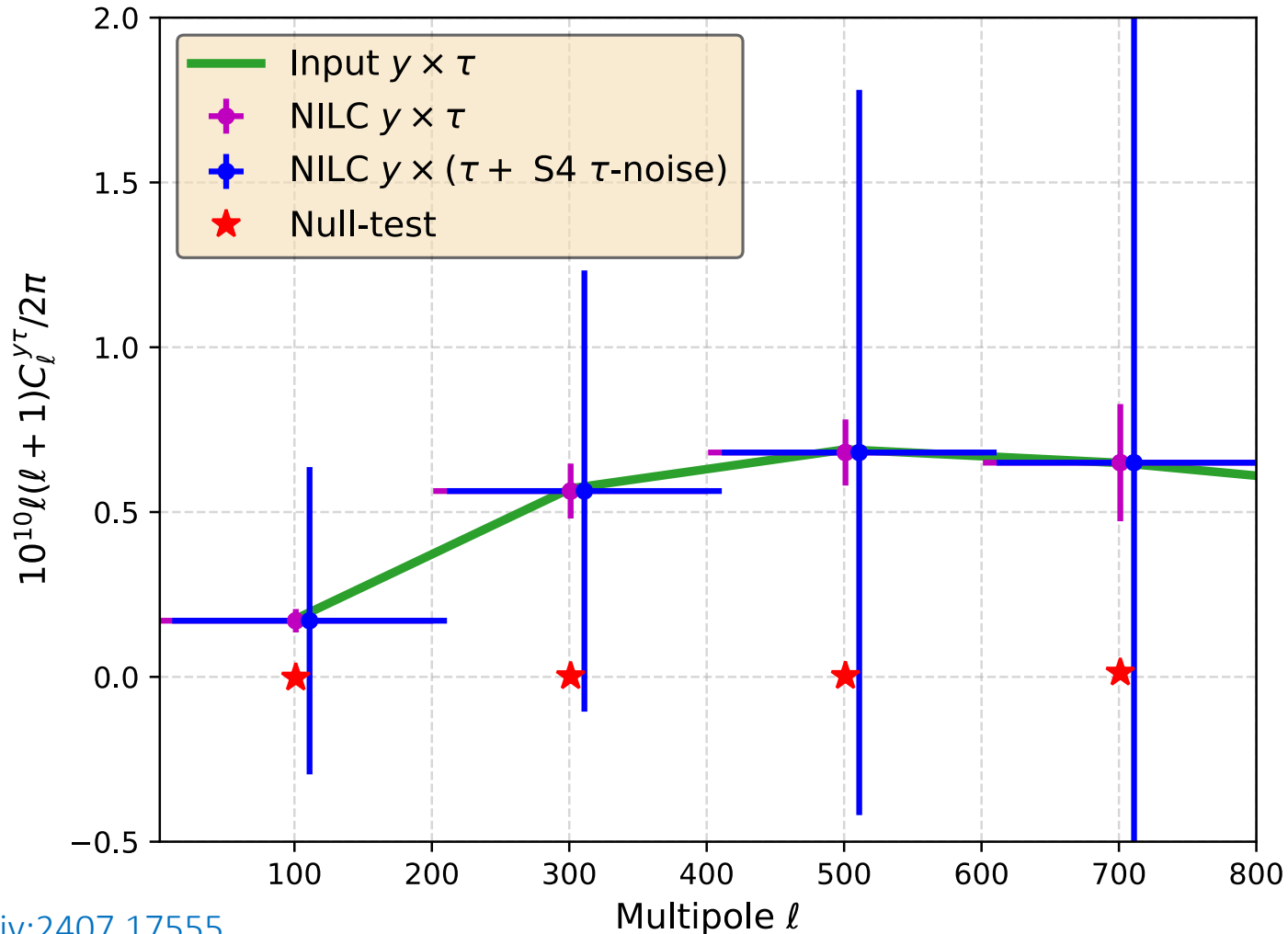
15% reduction in
uncertainty on
 $S_8 = \sigma_8(\Omega_m / 0.3)^{0.5}$
from the combined
LiteBIRD-Planck y-map



Thermal SZ effect from patchy reionisation

Cross-correlating the LiteBIRD SZ map with the CMB-S4 optical depth map

(following [Namikawa, Roy, Sherwin, Battaglia, Spergel, PRD 2021](#))



LiteBIRD will provide preliminary evidence of the faint thermal SZ signal from patchy reionisation with a modest SNR of 1.6



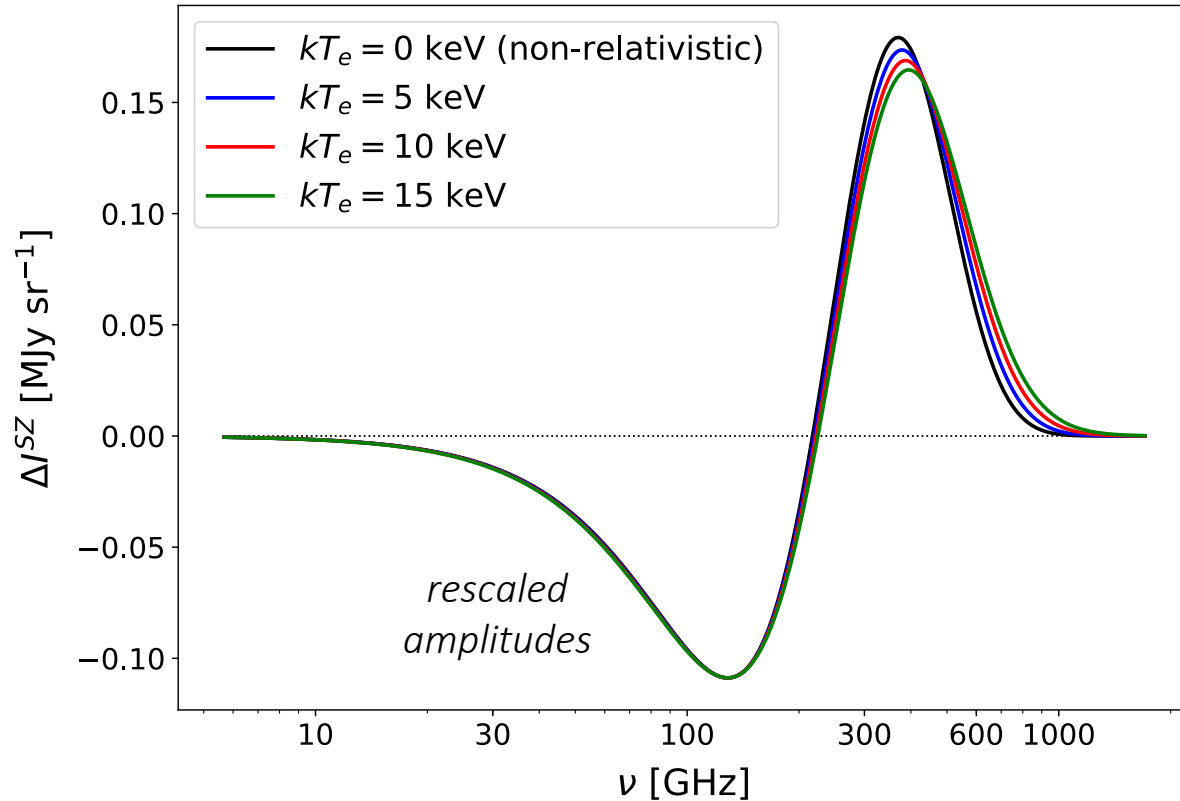
Perspectives on diffuse SZ science from a clean all-sky LiteBIRD y -map

- Relativistic SZ effect and gas temperature (capitalizing on LiteBIRD's high frequencies > 300 GHz)
- ISW-SZ cross-correlation at large angular scales
- CMB monopole y -distortion
- Two-halo contribution to SZ power spectrum at low multipoles
- Testing theories of structure formation via hot-gas tomography from SZ-LSS cross-correlations
- Quadrupole-like SZ effect from structures in local Universe such as the Milky Way or local supercluster
- SZ-coloured dipole-modulated CMB anisotropies via SZ-CMB cross-correlation as an alternative measurement of the dipole with higher significance than Planck Collaboration LVI (2020)
- Testing decaying dark matter models with SZ



Relativistic SZ effect

$$I_\nu(\hat{n}) = y(\hat{n})f(\nu, T_e(\hat{n}))$$

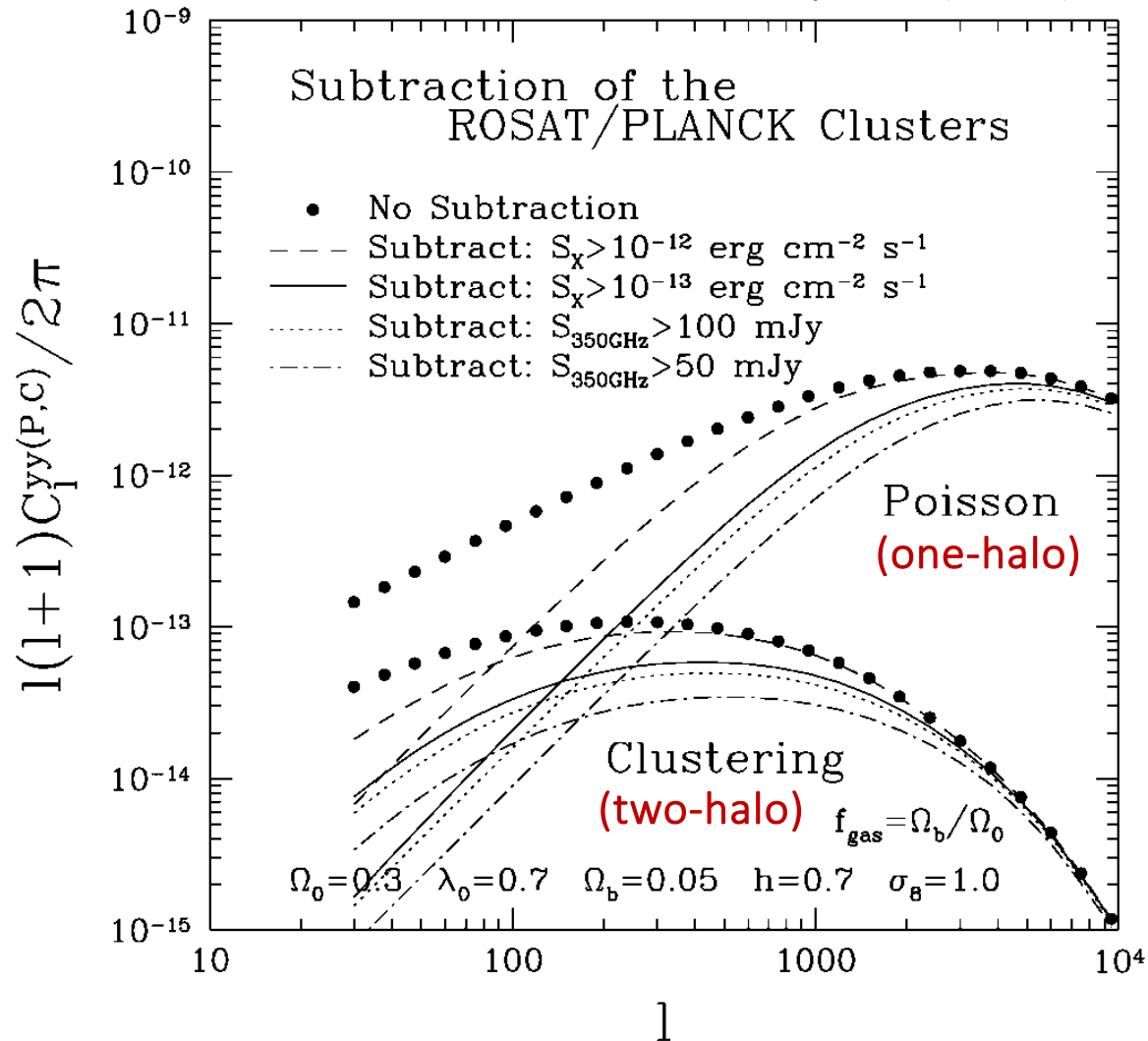


- *Capitalizing on LiteBIRD + Planck high frequencies above 300 GHz to disentangle the relativistic SZ effect*
- *LiteBIRD narrow bandpasses will also help detection*



Two-halo contribution to diffuse SZ effect

Komatsu & Kitayama (1999)

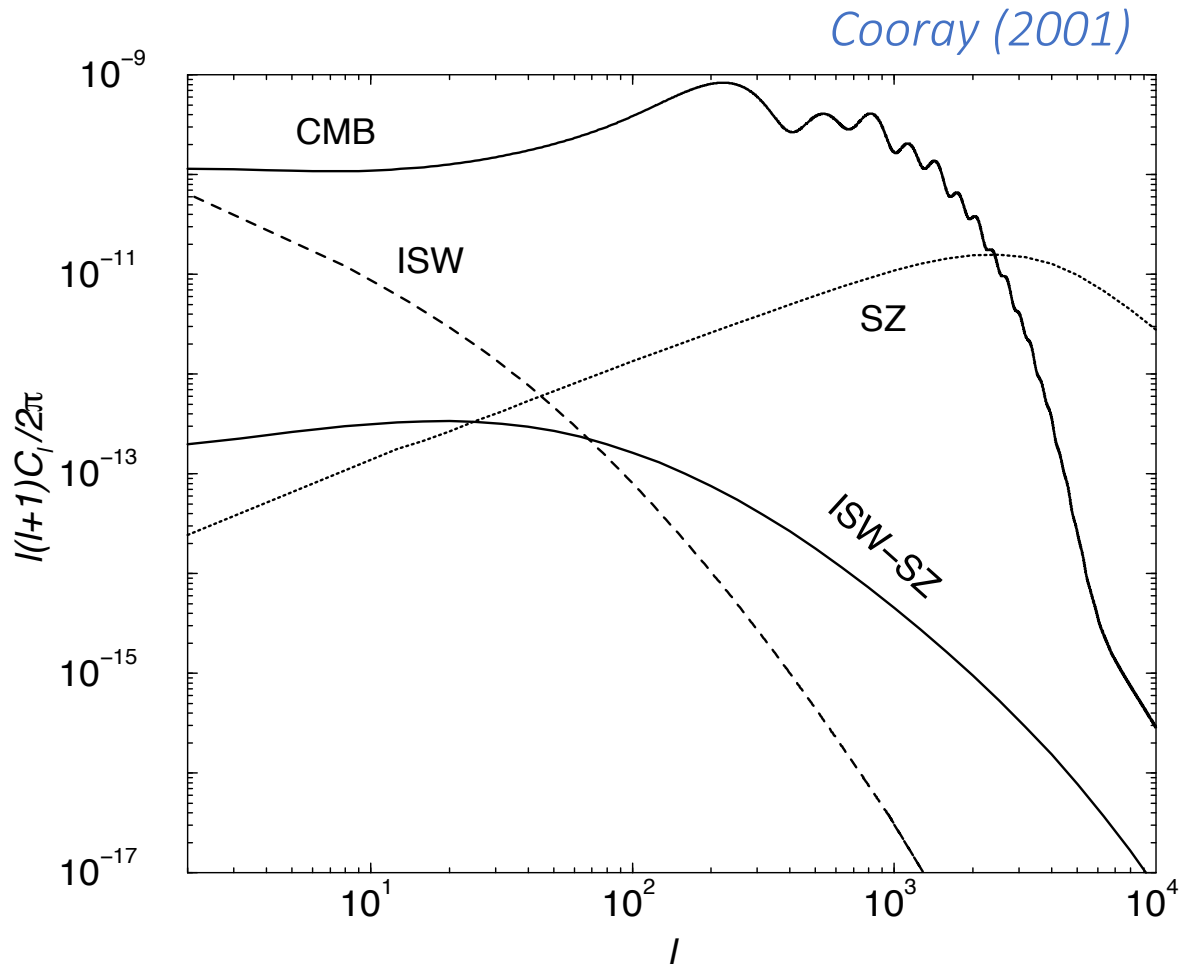


- Masking massive clusters in the y -map to “gaussianize” it and enhance the two-halo contribution over the one-halo signal in the diffuse SZ power spectrum at low multipoles

See “Rotti, Bolliet, Chluba, and Remazeilles, MNRAS (2021)”



ISW-SZ cross-correlation at large scales



- *Expected y - T cross-correlation at large angular scales between SZ and CMB temperature anisotropies due to the ISW effect*
- *LiteBIRD all-sky SZ map provides access to largest angular scales*



Conclusions

- An all-sky map of the thermal SZ Compton y -parameter probes the entire hot gas distribution across the Universe
- LiteBIRD's enhanced sensitivity and frequency coverage outperform Planck's SZ mapping results over the entire sky
- The combined LiteBIRD-Planck SZ map leverages both Planck's angular resolution and LiteBIRD's sensitivity
- Noise and foreground contamination reduced by a factor of 10 at large and intermediate scales in the combined LiteBIRD-Planck SZ map compared to the Planck SZ map
- Constraints on $S_8 = \sigma_8(\Omega_m/0.3)^{0.5}$ improved by 15% compared to Planck SZ map
- Many perspectives on diffuse SZ science from the all-sky LiteBIRD y -map